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Cretaceous accretionary complex related to Okhotsk-Chukotka Subduction, Omgon Range, Western Kamchatka, Russian Far East

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Abstract

The Omgon Range of Western Kamchatka contains a mid to Upper Cretaceous sequence of flysch with tectonic inclusions of Jurassic– Cretaceous oceanic rocks inferred to have been imbricated together in an accretionary prism. These rocks were tectonically juxtaposed during the Cretaceous in a mélange that contains a number of elements but mainly includes: (1) Middle Jurassic–Lower Cretaceous volcanic rocks formed in an oceanic and/or marginal sea environment; and (2) Albian–Campanian terrigenous turbidites made of quartz-rich clastic sediments that accumulated near a continental-margin. The oceanic rocks are inferred to have been tectonically incorporated into the continental terrigenous unit by offscraping during subduction. The accretionary prism resulted from subduction of the Pacific paleo-oceanic plate (Izanagi) under the Eurasian continental margin, which ultimately caused volcanism in the inboard Okhotsk-Chukotka volcanic belt. Internal imbrication was completed by the Maastrichtian (~70 Ma) as indicated by apatite fission-track ages that record cooling and exhumation of this crustal block. The Omgon accretionary wedge originated in a similar geodynamic setting and same time as the Yanranai (northern Korayk), Tonino–Aniva (southeastern Sakhalin), Hidaka (northeastern Japan) and Cretaceous part of the Shimanto belt (southwestern Japan). The similarities of ages, lithology, and tectonic setting suggest that the Omgon accretionary wedge was part of a paleosubduction zone along the Eurasian margin during the mid to Late Cretaceous.

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Keywords: Fission-track; Accretionary wedge; Mélange, Cretaceous; Kamchatka

1. Introduction

Northeast Asia includes oceanic and arc terranes of Paleozoic to Cenozoic age that have been swept into the margin from the Late Jurassic to Tertiary (Stavsky et al., 1990; Nokleberg et al., 1998). A dominant feature of the geology along this margin is the Cretaceous Okhotsk-Chukotka volcanic belt (OCVB), which is a continental arc built on part of this collage of terranes. The OCVB represents a laterally extensive Andean-style arc that persisted along the southern margin of the western and northern edge of what is now the modern limit of the Sea of Okhotsk and Bering shelf (Filatova, 1988; Nokleberg et al., 1998). The duration of magmatic activity in OCVB is

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debated but generally thought to include the Middle Albian to Campanian time (Belyy, 1977, 1994; Filatova, 1988; Zonenshain et al., 1990; Kotlyar et al., 2001). It sits on a collage of terranes that were assembled prior to the Albian. Calc-alkaline volcanic rocks from the exterior zone show that magmatism occurred locally between 86 and 81 Ma, and basaltic rocks with within-plate geochemical affinities yielded ages ranging from 78 to 74 Ma (Hourigan, 2003). The outboard forearc to this continental arc is less well defined and only partly exposed due to cover by adjacent offshore marine basins and younger strata.

One important question for regional tectonic reconstruc-tions is the continuity and lateral extent of subduction accretion associated with this important continental arc, and this question highlights two primary motivations for our focus on these units. First, we are interested in under-standing the history of long-term subduction accretion of oceanic material to the paleomargin. This material includes a number of oceanic complexes and turbiditic assemblages that record the dispersal of petrotectonic elements along

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the margin, mostly northward on Pacific plates. Second, we 113 are interested in understanding the continuity of petrotec-114 tonic assemblages along the paleomargin, which has been 115 disrupted and covered by younger elements. One of those 116 117 elements is normal faulting in the Sea of Okhotsk, which has 118 disrupted accreted assemblages through progressive north-119 to south-extension and wholesale crustal thinning. Another 120 is the deposition of volcanic and sedimentary sequences 121 related to subsequent arc-related volcanism, which is 122 represented by the modern Kamchatka-Kurile arc. Together, 123 the continuity of the Cretaceous paleomargin has been 124 obscured and therefore correlation of disparate blocks has 125 been challenging. 126

A detailed understanding of the lithology, structural 127 history, age and origin of internal blocks, and thermal 128 history of accretionary complexes can provide important 129 insight into the their setting and mode of emplacement. One 130 distinctive attribute of accretionary complexes is their 131 thermal history. Fission-track studies of exposed accre-132 tionary wedges have been carried out in a number of places 133 in the Pacific Rim including southwest Japan (Shimanto 134 belt; see Hasebe et al., 1993; Hasebe and Tagami, 2001), 135 and the west coast of North America (Franciscan Complex; 136 Dumitru, 1989; and Olympic Subduction Complex; Brandon and Vance, 1992; Brandon et al., 1998; Stewart 169 and Brandon, 2004). The accretionary wedge complexes 170 related to the OCVB subduction are poorly known because 171 much of the likely outcrop belt is offshore in the Bering Sea 172 and Sea of Okhotsk, or in a remote and logistically 173 complicated region on the northern Kamchatka Peninsula. 174 Therefore, little is know of these units, which we believe are 175 a crucial to paleotectonic reconstructions in this area. 176

Our research is focused on rocks of very limited 177 178 geographic extent in the Omgon Range, which are some 179 of the westernmost pre-Tertiary rocks on the Kamchatka 180 Peninsula (Fig. 1). In fact there are no other rocks like 181 this on the Kamchatka peninsula, so they provide a 182 glimpse of what is largely buried geology in this area. Our 183 study is based on mapping, structural analysis, geochem-184 istry of included blocks, and fission-track dating of detrital 185 zircon and apatite from sandstones of the flysch complex. 186 These studies resulted in the conclusion that the rock 187 complexes of the Omgon Range are fragments of a 188 Cretaceous accretionary prism presumably related to 189 OCVB subduction. As such, this small isolated exposure, 190 largely surrounded by Tertiary sedimentary strata, is 191 perhaps the single remaining onland exposure of this 192 tectonostratigraphic unit.



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225 **2. Geological overview**

The geology of the Kamchatka Peninsula is dominated 227 228 by the Neogene arc and related cover rocks. Pre-Tertiary rocks in Kamchatka generally occur as isolated exposures, 229 230 but are present throughout the Peninsula. Pre-Tertiary rocks in Western Kamchatka are uncommon and exposed as 231 widely scattered, isolated remote exposures that are very 232 difficult to access due to limited infrastructure. Although 233 234 pre-Tertiary rocks are exposed along the western coast of Kamchatka, rocks in the Omgon Range are quite distinct 235 236 from others in age and structure (Fig. 1), and have long been regarded as unique. 237

The Omgon Peninsula juts out into the poorly understood 238 Sea of Okhotsk, and therefore the rocks likely provide 239 240 insight into the evolution of this basin. The Sea of Okhotsk is floored almost exclusively by thinned continental crust. 241 Relatively little is known of the age and mechanism of 242 formation of the Sea of Okhotsk because the nature of 243 basement lithologies is almost completely unknown 244 (Hourigan, 2003). Crustal rocks that floor the Sea of 245 246 Okhotsk have been considered an example of a: (1) captured oceanic plateau (Watson and Fujita, 1985; Bogdanov and 247 Dobretsov, 2002); (2) accreted microcontinent block 248 (Parfenov and Natal'in, 1977; Zonenshain et al., 1990; 249 Konstantinovskaia, 2001), or (3) extended continental 250 framework of accreted terranes (Hourigan, 2003). Either 251 way, the lack of real data has allowed flexibility of 252 geological models. One of the biggest problems is that 253 basement rock has never been drilled, so most of what is 254 known is inferred from seismic data or dredging, but 255 dredged material is regarded as suspect due to possible ice-256 rafting. 257

The collided-block model holds that an allochthonous 258 sialic or oceanic plateau collided with the margin of 259 northeastern Asia in the Late Cretaceous, resulting in the 260 261 cessation of magmatism in the Andean-style Okhotsk-Chukotka belt (Parfenov and Natal'in, 1977; Zonenshain 262 et al., 1990; Konstantinovskaia, 2001; Bogdanov and 263 Dobretsov, 2002). In some models, the microcontinental 264 block forms the basement of the Sea of Okhotsk and extends 265 on land on the Kamchatka Peninsula. One such place where 266 this rock is exposed onshore may be in the metamorphic 267 rocks in the Sredinniy Range (Khanchuk, 1985). One new 268 269 hypothesis is that the Omgon-Palana belt (north and west of Sredinniy) is a collision zone separating the Sea of Okhotsk 270 plate from the West Kamchatka microplate (Bogdanov and 271 Chekhovich, 2002). They surmise that a fragment of an 272 ancient oceanic plateau makes up the Sea of Okhotsk 273 microplate, while the West Kamchatka plate is quasi-274 continental crust (Bogdanov and Chekhovich, 2002). A 275 back-arc extensional model for the origin of the Sea of 276 277 Okhotsk has recently been proposed, where basement rocks are inferred to be comprised of a variety of Mesozoic 278 279 terranes in a south-propagating extensional zone that has progressively extended since the Eocene (Hourigan, 2003). 280

Previous work has revealed that rocks of the Omgon 281 Range include an imbricated complex of volcanic, siliceous, 282 and carbonate rocks (Lower Cretaceous Kingiveem For-283 mation) and a terrigenous complex (Lower to Upper 284 Cretaceous Omgon Series) (Markovsky, 1989). The 285 volcanic-siliceous rock complex (Kingiveem Formation) 286 has been dated as Middle Jurassic (Bajocian-Bathonian) to 287 Lower Cretaceous (Lower Aptian) based on radiolarians 288 (Bondarenko and Sokolkov, 1990; Bogdanov et al., 1991; 289 Vishnevskaya et al., 1999). Flora and fauna from the Omgon 290 Series indicate that accumulation of the terrigenous rocks 291 occurred in the mid Cretaceous (in this case, Albian to 292 Santonian) (Vlasov, 1964). 293

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3. Geological structure of the Omgon Range

Following our mapping and field investigations, we 298 subdivide Mesozoic rocks of the Omgon Range (Fig. 2) into 299 a volcanic complex and a terrigenous complex, which are 300 tectonically interleaved. The volcanic complex consists of 301 pillow and massive aphyric, olivine-plagioclase and 302 plagioclase-microphyric, commonly amygdaloidal, basalt, 303 dolerite-basalt, and dolerite enclosing interlayers and lenses 304 of chert, siliceous mudstone, and, rarely, pelagic limestone. 305 All rocks of the volcanic complex make up fault-bounded 306 blocks and sedimentary slide blocks in the rocks of the 307 terrigenous complex. The terrigenous complex consists of 308 sandstones, siltstones, and mudstones, commonly with a 309 flysch-like alternation and interbedded thick conglomerate 310 beds. 311

Several tectonic slices have been mapped in the southern 312 segment of the Omgon Range. These slices consist of 313 volcanic rocks in panels that dip almost exclusively to the 314 northwest (Fig. 2). Structural observations at site 3 (Fig. 2) 315 reveal that both the volcanic and the terrigenous rocks dip 316 predominantly to the northwest (Fig. 3E) and faults that 317 bound the slides and blocks dip to the west (Fig. 3F). An 318 oblique relationship between the average fold axis (π -axis) 319 and fault strike probably reflects a strike-slip component of 320 displacement along a master fault. 321

While most folds in the Omgon rocks verge north and 322 northwest, there is generally a chaotic distribution of fold 323 axes (Figs. 2 and 3C,D). This observation may suggest that 324 the southern part of site 2 experienced rotation because 325 vergence differs dramatically from vergence of rocks at sites 326 3 and 1. Block rotation may be additional evidence for 327 strike-slip displacement. Two kilometers south of Cape 328 Promezhutochny, rocks of the terrigenous complex are 329 truncated by a nearly vertical, northeast-trending fault with 330 a strike-slip component (see Fig. 2), but the direction of 331 strike-slip movement is not clear. Terrigenous rocks do not 332 enclose any blocks of volcanic rocks north of this fault (site 333 1; see Fig. 2). Competent terrigenous rocks (sandstones and 334 conglomerates) make up a large southeast-northwest-335 trending anticline (Fig. 3A) and more plastic thin-bedded 336

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shale in the core of the anticline are isoclinally folded with 382 chaotically oriented fold axes (Fig. 3B). This pattern might 383 have been produced by deformation of poorly lithified 384 sediments or disharmonic folding. Numerous sills of 385 gabbro, diorite, quartz diorite, granodiorite, and leucocratic 386 granite, as well as quartz monzonite and granite-porphyry, 387 cut the rocks of the terrigenous complex at site 1 (Ledneva, 388 389 2001).

Non-marine, coal-bearing strata of the Middle Eocene
Snatol Formation unconformably overlies the deformed and
folded Mesozoic rocks (Gladenkov et al., 1991). This sharp

angular unconformity between the terrigenous complex and 438 the Eocene rocks has been described in the northern part of 439 the Omgon Range (site 2, Fig. 2). Here, basal conglomerates 440 consist of lithologies typical of the underlying pre-Tertiary 441 rocks of the Omgon Range (volcanic and terrigenous rocks) 442 and of the crosscutting sills. The Snatol Formation is folded 443 near the contact into tight to isoclinal folds with a northwest 444 vergence (Fig. 3G). These asymmetric folds suggest a local 445 displacement of the Eocene deposits northwestward 446 (Fig. 3H). The folding of the Tertiary deposits becomes 447 less intense with distance from the pre-Cenozoic rocks, and 448

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Fig. 3. Results of the structural-kinematic analysis of rock complexes in the Omgon Range (West Kamchatka). A-H are stereonets of various structural elements: A and B are for site 1 (Fig. 2): A, bedding planes; B, fold axes; C and D are for site 2 (Fig. 2): C, bedding planes; D, axial planes and axes of folds; E and F are for site 3 (Fig. 2): E, bedding planes; F, faults; G and H are for the Eocene deposits (Fig. 2): G, bedding planes; H, axes of asymmetric and symmetric folds. The linear and planar elements are shown with poles on a Schmidt net as projections on the lower hemisphere. N is the number of the structural elements of this type used for plotting the diagrams.

1.5 km east of the mouth of the Mainach River, these deposits are folded in a gentle east-dipping monocline (Fig. 2).

These structural observations indicate that rocks in the Omgon Range experienced at least two deformations. The younger deformation must be post-Middle Eocene, and it resulted in the folding of the Middle Eocene rocks (and the underlying sequences), with the principal axis of contraction oriented southeast-northwest (Fig. 3G,H). The older deformation must be pre-Eocene, as this is the age of rocks that rest above the unconformity. The tectonic interleaving of the terrigenous and volcanic rocks likely occurred during this first stage. The deformed rocks are Albian–Santonian, and because deformation may have been contemporaneous with deposition of the terrigenous rocks, we suspect that at least some of the earlier deformation occurred before Eocene deformation.

4. Composition and age of the rocks in the Omgon Range

4.1. Volcanic complex

The volcanic complex consists of sheets of pillow and massive, commonly amygdaloidal, basalt, ophitic dolerite, and dolerite enclosing interlayers and lenses of chert, siliceous mudstone, and minor limestone. Basalts at the top of sheets are represented by aphyric clinopyroxene-plagioclase and plagioclase-microphyric varieties. They display sheaf-like and, more rarely, hyalopilitic and vitrophyric textures in the groundmass. Glass is completely replaced by an aggregate of light-green chlorite and finely disseminated magnetite; there are a few spilitized rocks. The central and basal parts of the sheets consist of well-crystallized, medium-grained doleritic basalts and dolerites

composed of elongated and tabular plagioclase crystals and 561 idiomorphic and subhedral crystals of clinopyroxene and 562 magnetite, the magnetite locally forming accumulations. 563 Small plagioclase laths are locally enclosed in larger 564 clinopyroxene crystals. The interstices between these 565 crystals are filled with an aggregate of radiated-axial 566 chlorite and finely disseminated magnetite. The rocks 567 have a doleritic texture. Clinopyroxene crystals are fresh, 568 while plagioclase crystals are almost completely replaced 569 with saussurite, and the magnetite locally shows evidence of 570 oxidation. Amygdules in the basalts are filled with various 571 minerals, carbonate and carbonate + magnetite being most 572 common, and the association of carbonate+quartz or 573 alkaline chlorite is less common. Veinlets in fractured 574 rocks are filled with the same minerals and, more rarely, 575 with tremolite, which indicates a supply of carbonate, silica, 576 and alkalis. The discovery of exfoliation tuffs in the basalt 577 sheets, as well as the presence of interlayers and lenses of 578 chert, siliceous mudstone, and limestone, suggests that these 579 basalts erupted in a subaqueous environment. The volcanic 580 rocks have been weakly metamorphosed under low-581 582 temperature and low-pressure conditions.

The rocks examined showed high loss on ignition (LOI = 583 4.55–12.44%) (Table 1), as well as metamorphic alteration. 584 Both observations preclude using most of the mobile 585 elements for reconstructing the geodynamic conditions 586 that existed during emplacement of the volcanic rocks. 587 The concentrations of SiO₂, MgO, and Fe₂O₃ decline 588 drastically with increasing LOI, which suggests that these 589 components have been removed during alteration. However, 590 the variations in the Fe₂O₃ and MgO abundances as a 591 function of the LOI values are identical, implying that ratios 592 between these components are essentially unmodified. 593

Variations in the contents of major- and trace-elements 594 vary with the Mg# $(100 \times Mg/(Mg + Fe_{total}))$ and allow us to 595 distinguish two petrologic types of basalt: (a) poorly and 596 597 moderately differentiated basalts (MgO=9.12-7.29%) with normal iron content (Fe*/Mg=1.17-1.69); and (b) highly 598 differentiated basalts (MgO=4.12-4.28%) with an elevated 599 iron content (FeO*/MgO=2.07-3.19). The behavior of 600 major and trace elements in both of the recognized groups of 601 rocks is compatible with crystallization from a melt with 602 increasing differentiation of olivine+clinopyroxene+pla-603 gioclase and clinopyroxene + plagioclase + magnetite. 604

605 The Na₂O and K₂O contents and the high FeO*/MgO ratios suggest that the volcanic rocks of both types are close 606 to the tholeiite series. Despite differences in some 607 petrochemical-specific features, basalts with normal and 608 elevated iron contents belong to the same geochemical type. 609 Basalts of both types are highly depleted in LREE relative to 610 HREE ((La/Yb)_N = 0.37-0.86 and 0.42-0.65, respectively) 611 and have uniform Zr/Y (1.24-2.76 and 2.28-2.95) and 612 613 Zr/Sm (23.5–27.7 and 25.5–30.2) ratios (Table 1). These parameters, together with discrimination diagrams (Fig. 4), 614 place them close to N-MORB type basalts from oceanic 615 spreading centers (and/or from those of marginal seas). 616

Regionally these are similar to the Upper Jurassic–Lower617Cretaceous N-MORB exposed at Cape Povorotny (Taigonos618Peninsula) and in the Talovskie Mountains (northern619Koryak) (Khanchuk et al., 1990; Grigor'ev et al., 1995;620Sokolov et al., 2001; Silantyev et al., 2000).621

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4.2. Age of the volcanic complex

Previous work has confirmed the age of the volcanic 625 rocks in the Omgon Range as Middle Jurassic-Early 626 Cretaceous based on radiolarians from interbedded chert 627 (Bondarenko and Sokolkov, 1990; Bogdanov et al., 1991; 628 Vishnevskaya et al., 1999). We collected our own samples 629 of siliceous rocks and undertook our own radiolarian 630 analysis (Soloviev et al., 2001), and these indicate a Late 631 Jurassic-Early Cretaceous age for the host rocks (determi-632 nations by T.N. Palechek). Buchia from this unit are 633 Valanginian in age for siliceous rocks from the volcanic 634 complex (Buchia inflata (Lahusen) and Buchia sublaevis 635 (Keyserling) identified by V.A. Zakharov) (see Kirillova 636 and Kiriyanova, 2003) (Table 2). 637

4.3. Terrigenous complex

Sandstones of the terrigenous complex are poorly sorted, 641 angular graywackes. The sandstones are quartz-feldspar-642 and feldspar-quartz graywackes. Various volcanic rock 643 fragments and mudstone fragments are present among the 644 rock fragments. Basalt, andesite, and rhyodacite fragments 645 occur among the volcanic rock fragments, as well as 646 fragments of devitrified glass. Mudstone fragments are 647 especially common (up to 25%) among the sedimentary 648 rock clasts. Second in abundance are sedimentary rock 649 fragments consisting of a fine-grained, tuffaceous material. 650 Siliceous sedimentary rock fragments (i.e. chert) are 651 relatively rare (1-4%). The sediments also contain a 652 minor, but common occurrence of dispersed coalified 653 plant detritus. Fragments of metamorphic rock fragments 654 (quartzite and mica schist) are rare (less than 3%), although 655 they occur in all samples investigated. The overall 656 composition suggests derivation from a dissected volcanic 657 arc as discussed by Shapiro et al. (2001). 658

The specific chemical features of mudstones from the 659 flysch succession support the conclusion that the clastic 660 rocks were derived from a dissected continental volcanic 661 arc. The mudrocks are comparable, in terms of the 662 abundance of HFS lithophile elements, as well as 663 intermediate and heavy rare-earth elements (REE), with 664 the average post-Archean shale (PAAS), whose compo-665 sition is generally assumed to correspond to the composition 666 of the upper continental crust (Taylor and McLennan, 667 1985). However, compared to PAAS, they are poorer in 668 large-ion lithophile elements (LILE) and light REE 669 (Fig. 5a,b). At the same time, the lithophile element 670 spectra (high values of LILE/HFSE ratios, distinctly 671 pronounced Nb anomalies (Nb/Nb*=0.49-0.55) and Ta 672

Table 1

Tm

Yb

Lu

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Sample	O-1/98	O-2/98	O-7/98	O-9/98	O-12/98	O-14/98	O-25/98
Major oxides	(wt%)						
SiO ₂	47.31	49.91	43.63	43.88	44.53	48.12	46.66
TiO ₂	1.31	1.65	1.86	1.03	1.48	1.08	1.43
Al_2O_3	14.00	12.81	13.04	15.41	13.68	14.40	12.85
Fe ₂ O ₃	13.87	9.07	10.22	10.81	12.23	12.26	13.49
MnO	0.16	0.16	0.23	0.22	0.19	0.18	0.17
MgO	7.40	3.94	3.67	8.31	6.82	8.08	3.81
CaO	7.91	10.73	12.25	7.94	12.04	8.50	9.62
Na ₂ O	2.08	3.56	3.46	3.00	2.20	2.30	3.40
K_2O	0.30	0.09	0.36	0.39	0.32	0.57	0.48
P_2O_5	0.17	0.17	0.21	0.14	0.16	0.15	0.14
LOI	5.81	8.58	12.44	9.73	6.78	4.55	8.63
Sum	100.32	100.68	101.38	100.86	100.43	100.20	100.69
#Mg	51.39	46.28	41.57	60.38	52.50	56.64	35.88
Trace and rar	e-earth elements (pp	om) 42	47	40	50	40	50
SC	48	42	47	48	50 254	48	50 221
V Cu	312	3/1	3/0	200	354	289	331
Cr Cr	210	1/0	148	544	118	203	93
CO N:	49	120	39 67	05	30 101	40	54 76
INI Cu	97	129	70	1/1	101	117	70
Cu Zn	140	138	/0	123	107	155	72
ZII Ph	73	92	5 1	/4	80 6.6	12	01 12.5
Sr.	136	0.7	138	215	0.0	9.5	12.5
V	24	37	138	215	35	27	31
T 7r	2 4 66	88	124	54	81	53	70
Nh	1 41	1.90	124	0.99	1 57	1.00	1 23
Та	0.12	0.14		0.08	0.11	0.06	0.09
Ra	105	112	133	201	187	305	172
Hf	19	2 4	3.4	15	23	15	2.0
W	0.2	0.1	23	0.2	0.1	0.1	0.2
Ph	0.52	0.77	0.76	1.96	0.55	0.35	1.16
Th	0.09	0.18	0.20	0.09	0.11	0.06	0.12
U	0.03	0.12	0.59	0.05	0.18	0.03	0.60
La	1.86	2.41	3.72	2.48	2.14	1.34	2.07
Ce	6.30	8.09	11.45	7.01	7.36	4.72	6.88
Pr	1.11	1.43	1.98	1.11	1.36	0.88	1.26
Nd	6.32	8 43	11.17	5.92	8.37	5.46	7.13
Sm	2.39	3 20	4.11	2.18	3.20	2.25	2.76
Eu	0.89	1.11	1.32	0.86	1.17	0.88	1.02
Gd	3.29	4.50	5.57	2.89	4.57	3.36	3.96
Tb	0.59	0.85	0.98	0.53	0.85	0.61	0.75
Dv	4.07	5.81	6.65	3.48	5.76	4.18	5.17
Ho	0.94	1.36	1.55	0.78	1.35	1.00	1.18
Er	2.63	3.78	4.23	2.05	3.62	2 70	3 38

Notes. $Mg#=100 \times Mg^{2+}/(Mg^{2+}+Fe^{2+})$. The samples were crushed and powdered at the Laboratory of Miralogical and Fission-track analyses at the Geological Institute, RAS using the jaw crusher, vibrating cup mill and jasper mortar. The major- and trace-element contents were determined by X-ray fluorescence at the United Institute of Geology, Geophysics and Mineralogy, Siberian Branch of Russian Academy of Sciences (Novosibirsk, Russia) using standard procedures for controlling accuracy and reproducibility of the analysis. The trace elements were analyzed using an Inductively Coupled Plasma Mass Spectrometer (PerkinElmer/SciexElan 6100 DRC) at the Institute of Mineralogy, Geochemistry and Crystal Chemistry of Rare Earth Elements (Moscow, Russia). Sample preparation was done using a low-pressure HF digestion.

0.30

1.95

0.28

0.54

3.42

0.52

0.62

3.87

0.56

anomalies $(Ta/Ta^*=0.32-0.37))$ in multi-element dia-grams, where the mudstone compositions are normalized to the primitive mantle (Fig. 5c), are identical to those in the volcanic rocks of the calc-alkalic series. This result suggests that the mudstones of the terrigenous complex were mainly produced by the erosion of upper continental crust.

0.38

2.59

0.39

Deviations in the composition of mudstones from that of PAAS were most probably caused by the predominant contribution of volcanics from the active continental margin or from an ensialic island arc.

0.39

2.46

0.38

The composition of mudstones was apparently controlled by variable contributions from several sources, such as,

0.49

3.31

0.49

0.57

3.78

0.57

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Fig. 4. The Th-Hf-Ta (Wood, 1980) and Y-Nb-Zr (DePaolo and Wasserburg, 1976) discrimination plots demonstrating the similarity of the basalts we studied
 to N-MORB.

the upper continental crust, reworked sedimentary rocks, a young undifferentiated arc, a young differentiated arc, and exotic components. The high Th/U ratio (3.06-3.77), as well as Th/Sc and Th/Zr ratios, suggest that the sediments were not involved in significant recycling and also imply the insignificant effect of weathering on the composition of the mudstones. High Th/U ratios (3.06-3.77>3.0) and negative Eu anomalies (Eu/Eu*=0.72-0.97) suggest that continental crust made a contribution. However, low Th/Sc (0.35- $0.52 \ll 1$) and La/Sc (1.00–1.49 $\ll 4.0$) ratios and moderate La/Th ratios (2.844-2.88), in combination with the rather low Hf content (5.29-5.61 ppm), suggest a significant erosion of acid volcanic rocks in an active island arc or along an active continental margin (McLennan et al., 1993). The high Cr/Ni ratios (1.94–2.00), in combination with the elevated vanadium contents (194-257 ppm), indicate ero-sion of basic volcanic rocks (i.e. Garver and Scott, 1995).

Thus, mudstones of the terrigenous complex were presumably derived from volcanic rocks in an active island arc build on continental crust that was locally dissected. The likely source is compatible with, but not restricted to, the mid-Cretaceous Okhotsk-Chukotka volcanic belt.

4.4. Detrital zircon thermochronology

Cooling ages of detrital zircon from sandstones can provide minimum constraining ages for the time of deposition as well as information about the cooling of source terrains (Garver and Brandon, 1994; Garver et al., 1999, 2000; Soloviev et al., 2002a,b; Bernet and Garver, in press). Fauna and flora fossils from these rocks indicate Albian-Santonian deposition for the terrigenous complex (Vlasov, 1964). Detrital zircons were separated from eight samples of sandstones from the terrigenious complex, and from one from the unconformably overlying Eocene strata

Table 2	
Trace-element contents in shales and	sandstone of the Omgon Cape

Sample	OM-26/98	OM-36/98	O-28(4)/98
Trace and rar	e-earth elements (ppm	1)	
Sc	21	19	13
Ti	5689	5574	6051
V	257	194	96
Cr	116	71	22
Mn	374.02	433.32	376
Co	24	14	13
Ni	60	35	23
Cu	36	34	17
Zn	134	115	58
Rb	66.1	103.4	68.5
Sr	287	136	108
Y	32	36	24
Zr	175	208	171
Nb	11.18	14.75	18.7
Та	0.58	0.83	0.71
Ba	578	355	704
Hf	5.61	5.29	4.67
Pb	13.55	19.1	10.77
Th	7.20	10.00	6.22
U	2.35	2.65	1.50
La	20.76	28.40	22.50
Ce	55.06	66.32	43.60
Pr	7.84	8.61	6.04
Nd	22.70	28.45	25.07
Sm	5.09	6.29	4.37
Eu	1.71	1.53	1.40
Gd	5.70	6.68	4.19
Tb	0.76	0.97	0.61
Dy	4.96	5.54	4.38
Но	0.94	1.27	0.86
Er	2.79	3.83	2.15
Tm	0.51	0.55	0.39
Yb	2.54	3.62	2.15
Lu	0.44	0.59	0.33

Sample numbers OM-26/98 and OM-36/98 belong to shales; sample 895 number O-28(4)/98 belongs to sandstone. Also see notes to Table 1. 896

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Fig. 5. Multi-element diagrams and REE spectra for the terrigenous rocks of the Omgon Range. Normalized PAAS values are after (Taylor and McLennan, 1985); primitive mantle (PM) values, after Hofmann (1988) and chondrites, after Sun and McDonough (1989).

(Table 3). The ages of individual zircon grains were determined by the external-detector method, and 45-75 zircon grains were dated for each sample (Wagner and Van den Haute, 1992; Garver et al., 1999; see Supplementary material, Table 3, Fig. 6A,B). Our analysis of the distribution of fission-track ages from the terrigenious complex (not for Eocene deposits) determined that the samples have two or three component populations (P) of the following main age groups: P1: 114-80 Ma; P2: 187-142 Ma; and P3: c. 250 Ma. The presence of zircons of various ages in the sandstones, as well as the lack of secondary metamorphic minerals, suggest that the rocks were not heated above the zircon closure temperature after deposition (approximately 200-240 °C) (Brandon and Vance, 1992; Garver et al., 2005). Hence, it is likely that the ages of the populations reflect the cooling events of the rocks in the source area, and not subsequent heating events. The youngest population, P1, has an age range of $114.5 \pm 7.2 - 80.0 \pm 4.1$ Ma (Albian to earliest Campanian;

see Fig. 6d). Zircons from the younger population are mainly colorless idiomorphic crystals, which are character-istic, but not diagnostic, of first cycle zircons. These zircons most probably originated from volcanic activity synchro-nous with the flysch accumulation. It has been demonstrated in a number of papers that the age of the youngest zircon population is close to the age of the rock deposits, provided that volcanic activity occurred in the immediate vicinity of the sedimentary basin at the time of sedimentation (Garver et al., 1999, 2000; Shapiro et al., 2001; Soloviev et al., 2002b; Bernet and Garver, in press). Therefore, rocks of the terrigenous complex accumulated from Albian to earliest Campanian time assuming volcanism was active in the source area.

Samples of sandstones and mudstones were collected in the Rassoshina River Valley (east of the Omgon Range); these rocks are overlain by chert and pillow basalts. Because no fauna remains were discovered in the flysch in these exposures, these FT minimum ages provide the first

TICLE IN

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Table 3

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1009	Table 5	
1010	Summary of detrital zircon fission-track	

N sample Unit		Nt	Age range (Ma)	P1	P2	P3	P4
The Omgon	Range						
OM41	Eocene deposits	42	34–306	45.2±3.2 (39.2%)	68.3±13.0 (17.9%)	101.2±9.7 (35.7%)	293.0±60.7 (7.1%)
OM3	Terrigenous complex	75	56–209	80.0±4.1 (94.6%)	175.7±50.5 (5.4%)		
OM39	Terrigenous complex	74	62–193	85.3±4.2 (95.2%)	167.8±33.6 (4.8%)		
OM30	Terrigenous complex	46	66–254	90.6±9.0 (52.8%)	151.3±17.3 (47.2%)		
OM27	Terrigenous complex	75	67–275	99.8±5.8 (83.2%)	187.0±27.9 (16.8%)		
OM24	Terrigenous complex	75	74–365	102.0±18.9 (18.8%)	142.2±12.0 (68.1%)	248.2±28.8 (13.1%)	
OM22	Terrigenous complex	60	82–423	114.5±7.2 (69.8%)	-	237.1±25.3 (30.2%)	
The Rassos	hina River Valley						
OM48	Terrigenous complex	70	62–297	79.5±8.0 (30.0%)	108.0±12.3 (49.6%)	179.3±28.0 (20.5%)	
OM50	Terrigenous complex	65	61–264	77.7±6.6 (49.7%)	96.6±11.4 (46.0%)	198.3±64.8 (4.3%)	

Note. N_t = number of grains; percentage of grains calculated in a specific peak; Age for each population is in Ma, uncertainties cited at $\pm 1\sigma$. Zircons were 1031 dated using standard methods for FT dating using an external detector. Mounts were etched in a NaOH-KOH at 228 °C for 15 and 30 h and then irradiated at 1032 1088 Oregon State with a fluence of 2×10^{15} n/cm², along with zircon standards and dosimeter CN-5. Tracks were counted on an Olympus BX60 at $1600 \times$, and a ζ -1033 factor of 348.2 ± 11.02 was used. Fission-track ages were computed using the program Zetaage 4.7 (Brandon, 1996). To discriminate the populations by age, 1034 we used the program Binomfit 1.8 (Brandon, 1996). 1090

1035 constraints on their age (see Supplementary material, 1036 Table 3, Fig. 6D). The ages of the young population of 1037 zircons are 79.5 ± 8.0 and 77.7 ± 6.6 Ma. Note that the 1038 sampled flysch sections east of the Omgon Range appear to 1039 be somewhat younger than rocks of the terrigenous complex 1040 of the Omgon Range. 1041

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4.5. Fission-track dating of apatite 1043

1045 Fission-track dating of apatite from sedimentary rocks allows reconstruction of the thermal evolution of the 1046 sedimentary deposits after deposition because the annealing 1047 temperature of typical apatite is $\sim 110 \pm 5$ °C (Laslett et al., 1048 1987). FT dating of apatite from the sandstones of the 1049 terrigenous complex (Table 4) demonstrates that low-1050 temperature cooling occurred between 74 and 58 Ma. 1051 Apatite FT ages of 6 samples (OM3, OM22, OM24, 1052 1053 OM27, OM30, and OM39) are about 70 Ma, which suggests exhumation and cooling to ~ 100 °C (a depth of c. 4 km 1054 with a geothermal gradient of 25 °C/km) during the 1055 Maastrichtian. The apatite age from sample OM3 (57.7 \pm 1056 7.0 Ma) suggests reheating during a thermal episode 1057 associated with local intrusion of a sill (see Table 4). 1058

The Upper Cretaceous flysch deposits (Rassoshina River 1059 valley) experienced a different thermotectonic evolution, 1060 1061 because they have AFT cooling ages of c. 38 Ma. This young cooling event might have been associated with the 1062 transient thermal affects of the Eocene Kinkil volcanic belt 1063 (Gladenkov et al., 1991; Soloviev et al., 2002a). 1064

4.6. Cenozoic rocks of the Omgon Range

Numerous differentiated sills (Fig. 2) of basalt, basaltic andesite, andesite, dacite, and rhyolite and their holocrystalline equivalents intrude deposits of the terrigenous rock complex in the northern part of the Omgon Range (Ledneva 1097 et al., 2001). The sills were deformed along with the 1098 enclosing terrigenous deposits. The age of the sills was determined by the fission-track dating of apatite and zircon 1100 (see Table 4), and apparently the sills cooled, and therefore 1101 were likely to have been emplaced, in the Late Paleocene (63-60 Ma).

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1103 A sandstone sample from the basal horizons of the 1104 Eocene Snatol Formation that unconformity overlies the 1105 deformed Cretaceous rock (sample OM41) was collected for 1106 fission-track dating of zircon. The sample has four 1107 populations of cooling ages for detrital zircon (see 1108 Table 3, Fig. 6C). The youngest population of zircons 1109 from the sandstone is 45.2 ± 3.2 Ma (Middle Eocene), 1110 which is equivalent to the known stratigraphic age of the 1111 unit (Soloviev et al., 2001). 1112

5. Interpretation

A basic conclusion from our study is that Jura-1117 Cretaceous oceanic volcanics are tectonically mixed with 1118 mid-Cretaceous continental margin sediments in a structural 1119 complex formed during the latest Cretaceous. Volcanic 1120

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Fig. 6. Probability density plots (A, B and C with histograms) for representative fission-track grain-age distributions from the Omgon Range (Western Kamchatka). Thick lines show probability density distribution, and dashed lines show the best-fit peaks, as reported in Table 3. The fission-track minimum age corresponds to the age of the youngest peak. Plots were constructed according to Brandon (1996). Age is plotted on a logarithmic axis. The probability density scale is the same for both the density plots and the histograms. Density units are given relative to dZ=0.1, which corresponds to an interval on the age scale approximately equal to 10% of the age. Plot (D) of fission-track zircon results for sandstone from Western Kamchatka (Table 3). Circles—minimum ages (young population P1), triangles (P2), squares (P3), rhomb (P4)—older peak ages, respectively. Error bars show the 63% confidence intervals.

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1233	Table	4
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Apatite and zircon fission-track data

No. sample	Unit	Elev (m)	$ ho_{ m s}$	N _s	$ ho_{ m i}$	N_{i}	$ ho_{ m d}$	п	χ^2	Age	$-/+1\sigma$	$U (\text{ppm}) \pm 2\sigma$
Omgon ran	ge											
Zircon												
O98-27	Sill (gabbro)	10	68.5	1766	5.03	1296	0.27	30	99.4	62.5	-3.3 + 3.5	231.0 ± 17.8
O98-43	Sill (gabbro)	15	105.0	1116	8.42	891	2.92	17	100.0	63.0	-3.8 + 4.0	350.9 ± 32.7
Apatite												
098-27	Sill (gabbro)	10	4.51	222	1.12	551	2.93	15	100.0	60.9	-6.7 + 7.6	15.2 ± 1.4
OM3	Terrigenous	5	3.53	194	0.76	419	3.10	20	98	73.9	-8.5 + 9.6	9.8 ± 1.0
OM22	Terrigenous	7	4.02	101	0.96	241	3.08	15	49.8	66.5	-9.0+10.4	12.4 ± 1.7
OM24	Terrigenous	3	2.76	130	0.44	209	3.06	15	0.2	73.7*	-12.8 + 15.5	5.8 ± 0.8
OM27	Terrigenous	7	2.04	91	0.45	201	3.05	15	97.5	71.3	-8.3+9.2	5.9 ± 0.9
OM30	Terrigenous	150	3.93	168	0.86	367	3.03	15	80.8	71.5	-8.5 + 9.6	11.3 ± 1.3
OM39	Terrigenous	0	4.38	247	1.18	665	3.01	25	80.3	57.7	-6.2 + 7.0	15.6 ± 1.4
Rassoshina	River											
OM48	Terrigenous	150	4.77	282	1.39	822	2.98	26	0.0	37.6*	-6.1 + 7.3	18.6 ± 1.5
OM50	Terrigenous	150	7.78	166	1.64	349	2.96	15	0.0	38.0*	-7.8 + 9.8	22.0 ± 2.5

1250 *Note.* In this table ρ_s is the density (cm²) of spontaneous tracks (×10⁵) and N_s is the number of spontaneous tracks counted; ρ_i is the density (cm²) of induced 1251 1307 tracks ($\times 10^6$); and ρ_d is the density (cm²) of tracks on the fluence monitor ($\times 10^6$); n is the number of grains counted; and χ^2 is the Chi squared probability in 1252 1308 percent. Fission-track ages $(\pm 1\sigma)$ were calculated using the ζ -method, and ages were calculated using the computer program and equations in (Brandon, 1253 1996). The ζ -factors were 104.32 \pm 3.35 (for apatite based on CN1 calibration) and 348.2 \pm 11.02 (for zircon based on CN5 calibration). All ages that pass χ^2 1309 1254 (>5%) are reported as pooled ages, otherwise first population ages calculated by BinomFit 1.8 (Brandon, 1996; Brandon, 2002) are shown (denoted by *). 1310 Glass (CN-1) monitors, placed at the top and bottom of all irradiation packages (for ζ calculations) were used to determine the fleunce gradient in each package. 1255 1311 After etching, mounts were covered with a low-uranium mica detector, and irradiated with thermal neutrons at Oregon State University with a nominal fluences 1312 1256 of 8×10^{15} n/cm² (for apatite) and 2×10^{15} n/cm² (for zircon), along with a standards (Fish Canyon Tuff, Buluk Tuff) and a reference glass dosimeter CN1 (for 1257 1313 apatite) and CN5 (for zircon). All samples were counted at $1600 \times$ using a dry $100 \times$ objective (10 oculars and $1.6 \times$ multiplication factor) on Olympus BX60 microscope fitted with an automated stage and a Calcomp digitizing tablet. 1258 1314 1259 1315

rocks of the Omgon Range formed during the latest
Jurassic–Early Cretaceous in an oceanic or marginal sea
setting. Basalts of this complex are comparable with
N-MORB from oceanic-type spreading centers. It is
possible that the paleo-Pacific-Izanagi plate (Engebretson
et al., 1985) was the source of the volcanic blocks.

Terrigenous rocks accumulated as turbidites in sub-1266 marine fans during the Albian to the Campanian in a 1267 continental-margin environment. The composition of the 1268 mudstones and sandstones suggest the source was a 1269 dissected volcanic arc, probably the Okhotsk-Chukotka 1270 volcanic belt, which was built on continental basement of 1271 the Eurasian margin. Blocks and slides of the volcanic rocks 1272 have tectonic contacts with the terrigenous rocks, which 1273 make up the matrix of the succession. 1274

Thus, rocks of different ages that were formed in 1275 different geodynamic settings are tectonically mixed. This 1276 mixing of oceanic lithologies within a matrix of terrigenous 1277 rocks suggests that the rock units of the Omgon Range are 1278 part of an accretionary prism. In this scenario, slides and 1279 blocks of oceanic origin were accreted during subduction 1280 and mixed with the terrigenous Albian-earliest Campanian 1281 deposits of the continental-margin. 1282

Fission-track dating of apatite suggests that this accretionary prism was exhumed to a near-surface level (<c. 4 km) by the Maastrichtian (\sim 70 Ma), about 10– 20 Myr after deposition. FT ages of zircon and apatite from felsic sills shed light on the level of exhumation of the accretionary complex in the Late Cretaceous. Most of the cooling ages (both ZFT and AFT) fall between 60 and 1316 70 Ma (Table 4), and some ZFT and AFT ages are nearly 1317 concordant (i.e. sample 98-28), which suggests that at that 1318 time, the enclosing rocks were at relatively shallow levels 1319 (<4 km). Intrusion of the felsic dikes and sills probably 1320 marked the end of accretion of material in this system. If this 1321 is the case, the accretion process had been completed by the 1322 Late Cretaceous, and rocks of the Omgon Range were 1323 incorporated into the structure of the continental margin. In 1324 the Late Paleocene, sills and dikes intruded into the 1325 accretionary prism at a latitude close to the present-day 1326 position of the Omgon Range as indicated by paleomagentic 1327 studies (Chernov and Kovalenko, 2001). This intrusion 1328 signals an extremely important oceanward shift in the locus 1329 of arc magmatism in the latest Cretaceous to Early Tertiary. 1330

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6. Regional setting

An important aspect of paleogeographic reconstructions 1335 is the regional distribution of arc and forearc complexes 1336 along the NE Eurasian margin. Previous studies have not 1337 fully investigated the lateral continuity of Cretaceous 1338 accretionary complexes along the northeastern Eurasian 1339 margin. This lack of analysis is partly due to discontinuity of 1340 exposures in remote, difficult to access locations. Because 1341 exposures are limited, there is considerable uncertainty in 1342 tectonic reconstructions for the evolution of the northeast 1343 Eurasian margin in the Cretaceous. Here we try to fill this 1344

1345 gap by piecing together relicts of the Cretaceous accre-1346 tionary complexes along the NW Eurasian margin.

Throughout much of the northwestern Pacific, accumu-1347 lation of terrigenous strata commenced in the Albian in 1348 basins genetically associated with subduction under the 1349 Eastern-Asian volcanic belt (Belyy, 1977; Filatova, 1988; 1350 Zonenshain et al., 1990; Belyy, 1994; Hourigan and Akinin, 1351 2004). The Eastern-Asian volcanic belt is a laterally 1352 extensive Andean-style arc subdivided into different sectors 1353 based on differences in the basement rock types and 1354 lithologic similarity of volcanic sections within specific 1355 1356 geographic regions (Belyy, 1977; Melankholina, 2000) (Fig. 7). From north to south these sub-divisions include: 1357 Chukotka-Alaska, Okhotsk-Chukotka, Eastern Sikhote-1358 Alin, and Korea-Japan. Terrigenous sedimentation occurred 1359 1360 in syn-subduction forearc basins along the eastern Eurasian margin (Melankholina, 2000; Garver et al., 2000; Shapiro 1361 et al., 2001; Soloviev et al., 2001; Zharov, 2003). These 1362 basins include, from north to south, the Bering Sea basin, 1363 Ukelayat basin, North Okhotsk basin, Western Kamchatka 1364 1365 basin, Western Sakhalin basin, Ieso basin, and the Shimanto 1366 basin. Clastic rocks in these basins accumulated in slightly different tectonic settings, but all record erosion of the arc 1367 and deposition along the continental margin. Continental 1368 and shallow-water forearc basins, and post-Albian molasse 1369 deposits, are known in the Pendzhina and Northern Korayk 1370 regions (Zinkevich, 1981; Filatova, 1988; Sokolov, 1992). 1371 The Albian-Campanian forearc basin deposits have been 1372 described in the Penzhina Guba (Bay) area northwest of 1373 1374 northern Kamchatka (Tuchkova et al., 2003). The Western Sakhalin and Ieso forearc basins (Fig. 7) contain Aptian-1375 1376 Paleocene clastic deposits derived from the Eastern Sikhote-Alin volcanic belt (Melankholina, 2000; Zharov, 2003). 1377

The continent-derived flysch with exotic ocean-derived 1378 blocks is typical and are interpreted to have been 1379 imbricated in accretionary wedges along the continental 1380 1381 margin (i.e. Cowan, 1985). The relicts of the accretionary wedge related to Cretaceous subduction are known in the 1382 following segments of the present-day tectonic framework 1383 of the eastern Eurasian margin (Fig. 7): Yanranai (Korayk 1384 Upland) (Grigor'ev et al., 1987; Sokolov, 1992), Omgon 1385 (Western Kamchatka) (this study), Tonino-Aniva (south-1386 eastern Sakhalin) (Zharov, 2003), Hidaka (Kiminami 1387 et al., 1992; Zharov, 2003) and Shimanto (Taira et al., 1388 1389 1988). Hence, we are impressed not only by the similarity of these units, but also the apparent lateral continuity of 1390 this N-S petrotectonic assemblage. We review the 1391 occurrences of these rocks below, starting in the north 1392 and working southward. 1393

The three tectonic slices were described in the Yanranai accretionary complex in the Koryak upland of northern Kamchatka (from Grigor'ev et al., 1987; Sokolov, 1992). The slices consist of oceanic crust fragments with different ages, but younger rocks have a lower structural position. The terrigenous rocks, inferred to have been derived from the Eurasian margin, occur in the upper part of the each



Fig. 7. Tectonic elements of the mid- to Late-Cretaceous margin within the 1442 context of the modern setting of northeastern eurasia (Melankholina, 2000). 1443 Outlined letters (Squares) represent the following sectors of the mid- to Late Cretaceous Eastern-Asian volcanic belt: a, Chukotka-Alaska; b, 1444 Okhotsk-Chukotka; c, Eastern Sikhote-Alin; d, Korea-Japan. Numbers in 1445 circles are syn-subduction basins: I, Bering Sea; II, Ukelayat; III, North 1446 Okhotsk basin; IV, Western Kamchatka basin; V, Western Sakhalin basin; 1447 VI, Ieso basin; VII, Shimanto basin. Letters in rhombohedra are fragments 1448 of the accretionary wedge: A, Yanranai; B, Omgon; C, Tonino-Aniva; D, Hidaka; F, Shimanto. 1449

slice. The accretion of the Jurassic–Neocomian oceanic 1451 rocks, presumably part of the Izanagi or Kula plate, 1452 occurred at the end of the Early Cretaceous. The second 1453 stage of accretion was at the end of the Late Cretaceous. The 1454 final stage was completed in the Maastrichtian when an 0listostrome unit formed. 1451

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Rocks of the Tonino-Aniva Peninsula (southeastern 1457 Sakhalin Island, western part of the Sea of Okhotsk) are 1458 inferred to have accumulated in an accretionary prism 1459 (Zharov, 2003). The Tonino-Aniva complex consists of a 1460 mid-Cretaceous turbidite and olistostrome unit with tectonic 1461 slivers of volcanic rocks inferred to be fragments of a 1462 Jurassic-Lower Cretaceous seamount. The structural lower 1463 part of the complex is represented by Upper Cretaceous 1464 turbidites. The Tinino-Aniva complex is inferred to have 1465 formed as a result of Aptian-Cenomanian subduction of the 1466 Sorachi oceanic plateau and syn-accretionary turbidites in 1467 1468 the Late Cretaceous.

Farther south, the Hidaka terrane on Hokkaido Island 1469 (Japan), has been described as a Late Cretaceous-Early 1470 Eocene accretionary wedge (Kiminami et al., 1992; Zharov, 1471 1472 2003). The terrane contains continental-derived terrigenous mélange that is probably a dismembered turbidite complex 1473 with an eastern structural vergence. The change from 1474 hemipelagic rocks low in the section to clastic rocks higher 1475 in the sequence shows that the depositional setting varied 1476 1477 from abyssal plain to continental margin.

1478 The best-studied example of the accretionary wedge from the entire belt is the Shimanto (Taira et al., 1988; 1479 Suzuki, 1988; Matsumoto et al., 1988; Hasebe et al., 1993; 1480 Hashimoto and Kimura, 1999; Hasebe and Tagami, 2001). 1481 The Shimanto Belt has an overall younging trend from north 1482 to south (Taira et al., 1988), but here we are mainly 1483 concerned with the older part of the belt that is similar in age 1484 to those to the north. The Upper Cretaceous part of the 1485 Shimanto Belt includes turbiditic sandstone and shale and 1486 minor conglomerate interbedded with hemipelagic varico-1487 lored shale. The biostratigraphic ages from the flysch unit 1488 range from Coniacian to Campanian. Mélange, which 1489 occurs as several linear belts sandwiched between flysch 1490 units, is composed of a highly deformed argillaceous 1491 'matrix' with various-sized tectonic slivers. The tectonic 1492 1493 slivers include pillow basalts, chert, and varicolored shale. A remarkably constant age-lithology relationship occurs in 1494 this unit: the tectonic slivers contain dated rocks from the 1495 Tithonian to Cenomanian, and the sheared argillaceous 1496 matrix yields mostly Campanian radiolaria (Taira et al., 1497 1988). Most of the meta-basalts in the Shimanto mélange 1498 zones are considered to have originally been ocean-floor 1499 basalts (MORB), but also include alkali basalts possibly 1500 1501 derived from volcanic islands or seamounts (Suzuki, 1988). The overwhelming geological evidence suggests that the 1502 Shimanto Belt is an accretionary prism of Cretaceous to 1503 Tertiary age (Taira et al., 1988). The contemporaneous 1504 Cretaceous volcano-plutonic belt is distributed along a 1505 linear belt extending from Japan to Shikhote Alin to the 1506 north. This same plutonic belt can be traced northward into 1507 the Okhotsk-Chukotka volcanic belt (OCVB). 1508

Our new data suggest that the Omgon accretionary
complex belongs to this family of mid to Upper Cretaceous
accretionary complexes that accumulated along the NW
Pacific margin (Fig. 7). The Omgon is similar to

the Yanranai (northern Koryak), Tonino-Aniva (south-1513 eastern Sakhalin), Hidaka (northeastern Japan) and Cretac-1514 eous Shimanto belt (southwestern Japan) in age, structure 1515 and tectonic position. Together, these accretionary 1516 complexes record Cretaceous subduction under northeastern 1517 Eurasian, assuming there has not been significant tectonic 1518 translation along the margin. One important aspect to note 1519 about this family of accretionary complexes is that the ones 1520 farthest south (Japan) appear to have continued from the 1521 Cretaceous to Eocene, but to the north there is no evidence 1522 of this complex having ages younger than the Cretaceous. It 1523 is possible that this difference can be attributed to accretion 1524 and outboard jump in accretion after the Cretaceous in the 1525 northern areas. 1526

The Omgon accretionary complex separated from the 1527 Okhotsk-Chukotka volcanic belt due to extension of the Sea 1528 of Okhotsk, the evolution of which is poorly known (see 1529 geological overview). We speculate that our new data are 1530 consistent with the back-arc extensional model of the origin 1531 of the Sea of Okhotsk (Hourigan, 2003). The basement 1532 rocks of the Sea of Okhotsk are inferred to be comprised of a 1533 variety of terranes that were rifted apart since the Eocene. 1534 One terrane is the Omgon accretionary complex and 1535 presumably an inboard forearc basin. 1536

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7. Conclusions

The Omgon Range in Western Kamchatka is composed 1541 of southeast-verging interleaved tectonic units imbricated in 1542 a subduction setting. Sandstones are uniform in composition 1543 and the sediment is inferred to have been derived from a 1544 continental arc. FT depositional ages of detrital zircons from 1545 the Omgon flysch are Albian-Campanian, which is similar 1546 to the depositional age inferred from fossils. It is likely that 1547 this arc that supplied terrigenous sediments was the 1548 contemporaneous Okhotsk-Chukotka volcanic belt partly 1549 because the main phase of volcanism and plutonism in this 1550 belt occurred during the mid-Cretaceous. The flysch is 1551 clearly imbricated with older oceanic rocks. The basalts are 1552 tholeites similar to those associated with spreading centers 1553 within oceanic and marginal basins, and the overlying 1554 siliceous rocks are Upper Jurassic to Lower Cretaceous in 1555 age. The Albian-Campanian continent-derived flysch with 1556 the exotic Upper Jurassic to Lower Cretaceous ocean-1557 derived blocks is a relict of the accretionary wedge related 1558 to Cretaceous subduction under the Eurasian margin. 1559 Internal imbrication of the Omgon complex was complete 1560 by the Maastrichtian (\sim 70 Ma) when apatite fission-track 1561 cooling ages were recorded. 1562

Other accretionary complexes to the north (northern 1563 Kamchatka) and to the south (Japan) consist of a broadly 1564 coeval terrigenous matrix and oceanic blocks, the blocks 1565 being older than the matrix. We suggest that these disparate 1566 occurrences of mid to Upper Cretaceous accretionary 1567 complexes retain a record and the gross position of 1568

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the subduction zone along the Eurasian margin. Cretaceous
to Eocene ages of matrix rocks to the south may indicate that
the subduction zone operated continuously to the south, but
was interrupted to the north. We speculate that the Omgon
accretionary complex was separated from the OkhotskChukotka volcanic belt by extension that occurred during
formation of the Sea of Okhotsk since Eocene time.

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1578 8. Uncited reference

Shutov et al. (1972).

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Appendix. Supplementary material

Supplementary data associated with this article can be found, in the online version, at doi:10.1016/j.jseaes.2005. 04.009

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